

# Image Calibration and Deconvolution for Space Object Imaging

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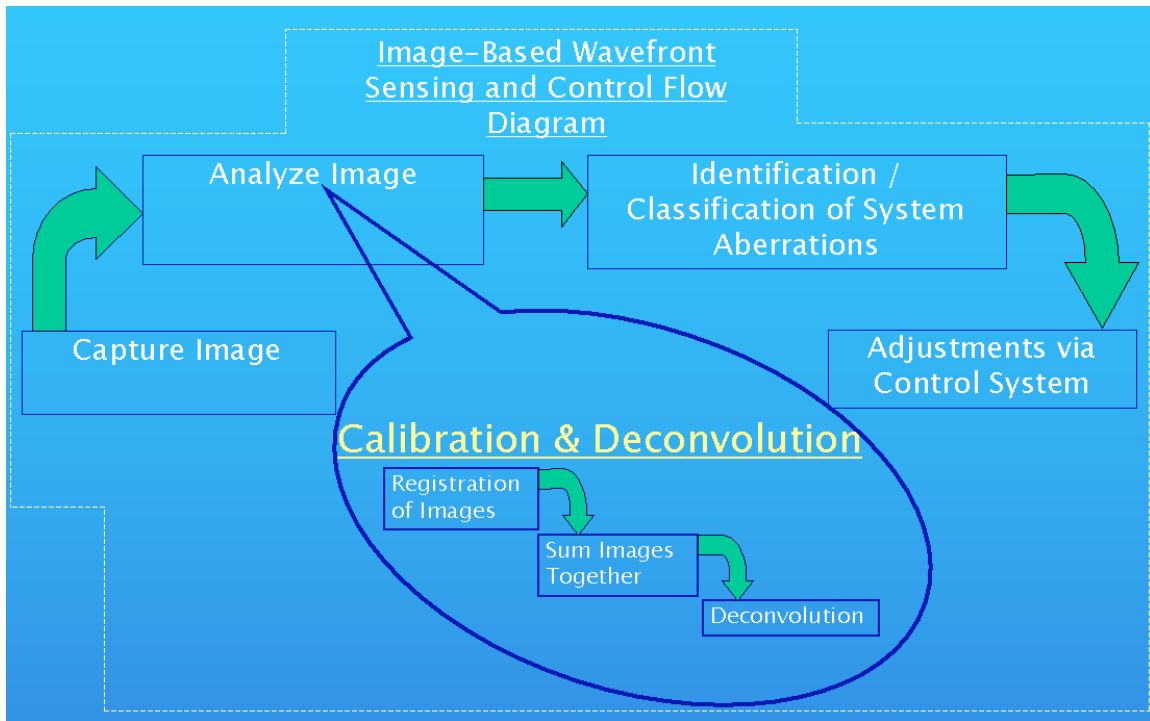
July 17, 2003

## Abstract:

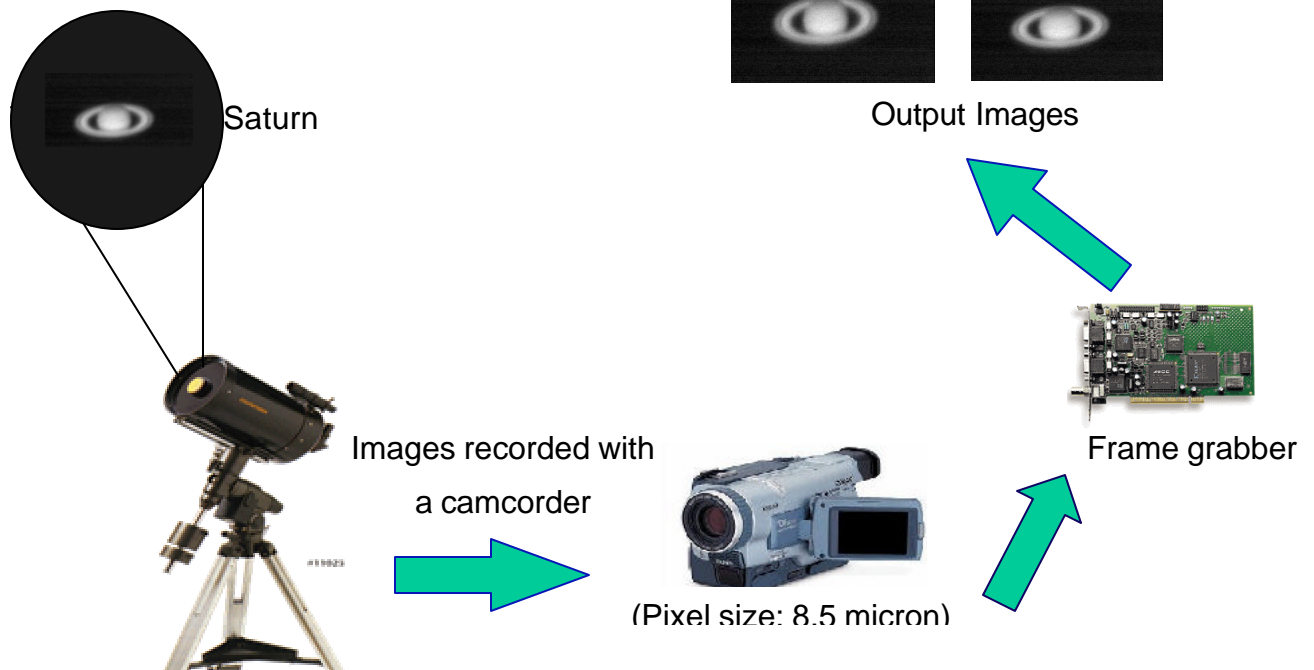
Data was collected using a 12" Schmidt-Cassegrain telescope with off-the-shelf imaging equipment. The data was then recorded using a home video recorder and the individual frames were captured from mini-DV tape. Video and CCD imaging systems can produce short-exposure images that contain little or no information due to poor SNR (signal-noise-ratio), blurring due to object motion, atmospheric turbulence, or optical system mis-alignments. Each of the listed aberrations can be corrected to some degree and faults within the optical system can be estimated using image calibration and deconvolution techniques.

## Introduction

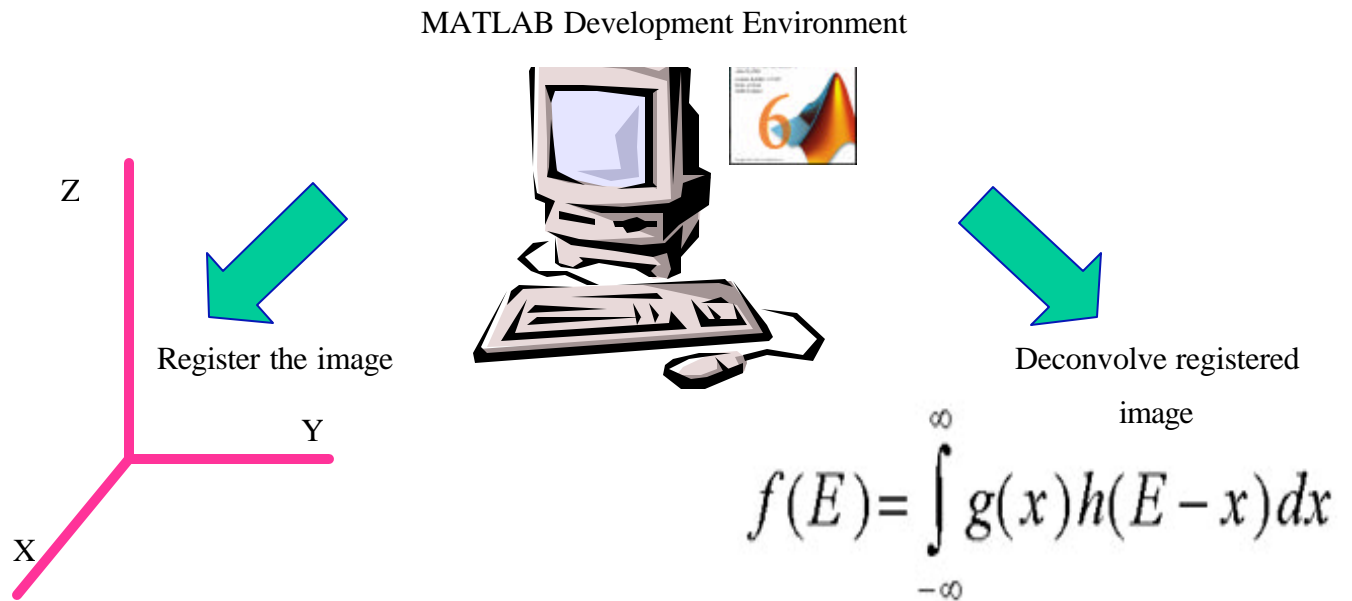
Testing the optical system can be achieved via analysis and repair of resultant images by means of calibration and deconvolution algorithms/filters. This can be used as one approach to Image-based wavefront sensing and control.



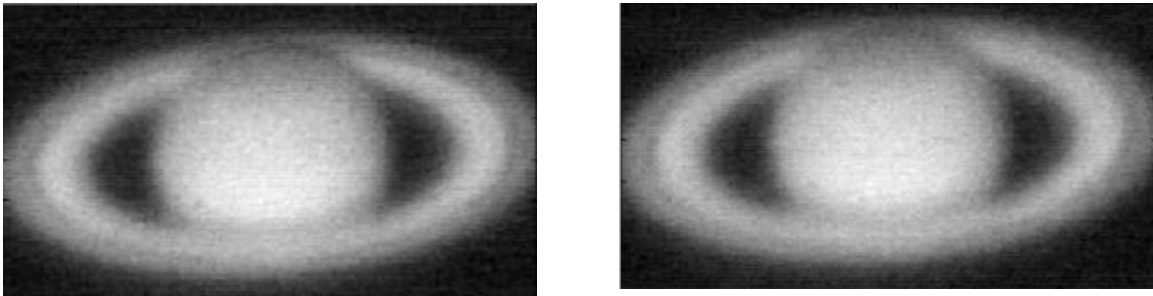
## Experimental Setup



## Analysis Methods and Tools



## Sample Data Frames

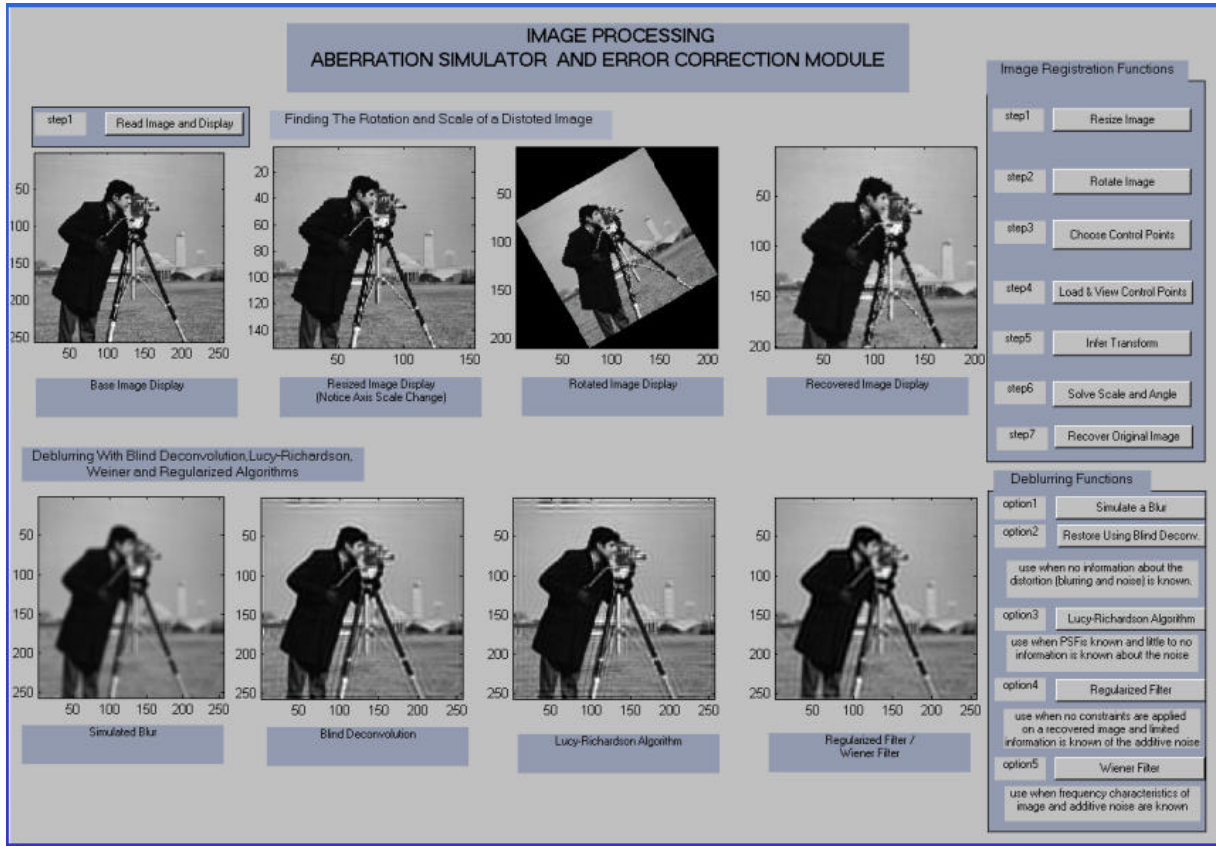


## Personal Contribution

Development/Testing of Calibration and Deconvolution Techniques

Development of a Graphical User Interface for Image Processing

## Graphical User Interface



## General Filter Derivation

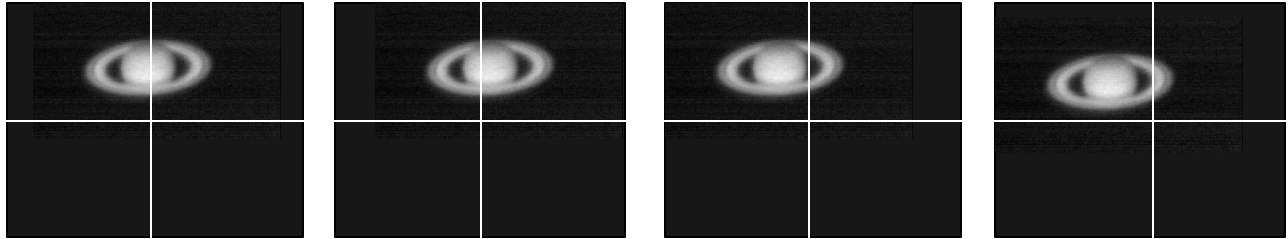
$$i_d = psf_A \otimes i_o + n, \quad \mathfrak{F}(i_d) = I_d = PSF \cdot I_o + N, \quad I_o' = Y \cdot I_d$$

$$\text{minimize: } e^2 = \left\langle \left| I_o - I_o' \right|^2 \right\rangle, \quad \frac{\partial}{\partial Y} (e^2) = 0, \quad \Rightarrow Y = \frac{PSF}{PSF^2 + (S_N / S_o)}$$

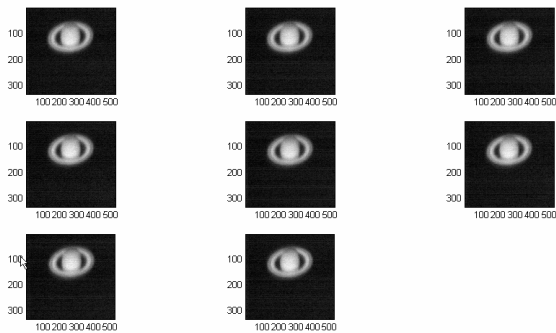
$$\text{Case1: } \frac{S_N}{S_o} \ll 1, I_o = \frac{I_d}{PSF} \quad \text{Case2: } \frac{S_N}{S_o} \gg 1, Y \rightarrow 0$$

## Results and Conclusions

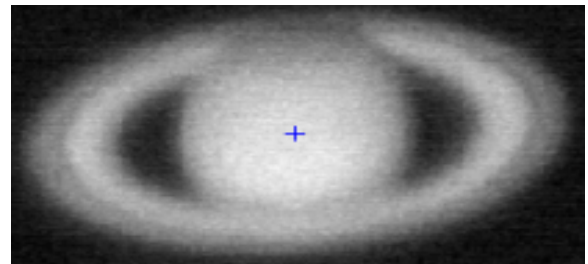
Sample Raw Data From Frame Grabber:



Calibrated Images:

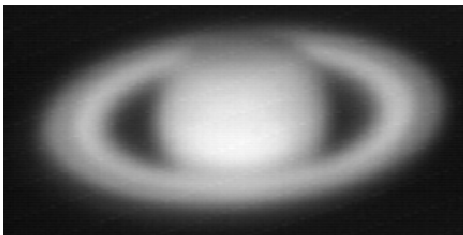


(Registration by hand)

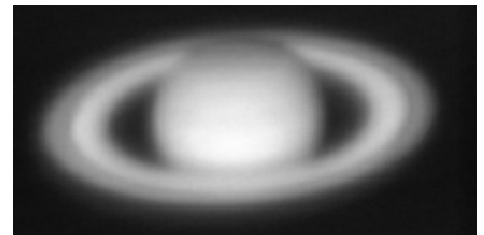


(Registration via centroiding)

Summed Images:

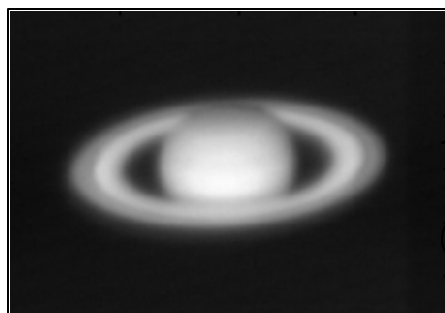


(No centroiding)



(Centroid ( tol:0.05) +Sum)

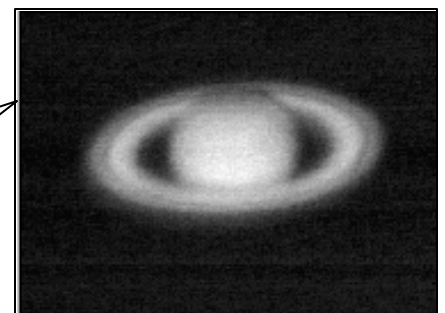
Deconvolved Summed Image:



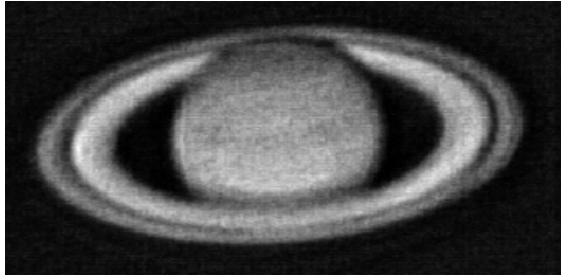
MATLAB Blind Deconvolution

Recovered

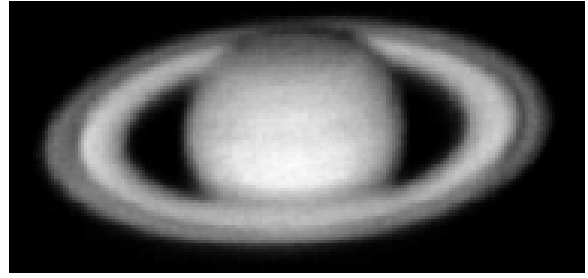
Single Frame



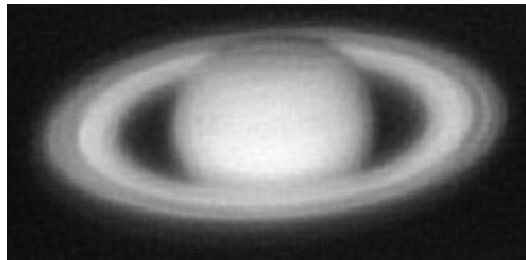
Alternative Approaches:



1



2



3

- Register, Deconvolve and Sum (1&2): Custom Ayers-Dainty Algorithm
- MATLAB: Deconvolve, Register and Sum (3)

In conclusion, calibration and deconvolution techniques can greatly improve image quality. In addition, resulting information can be used to characterize the optical system.